Physica D 71 (1994) 1-17 North-Holland

SSDI: 0167-2789(93)E0246-8



Self consistent chaos in the beam plasme instability

J.L. Tennyson^{a,1}, J.D. Meiss^b and P.J. Morrison^c

* Stanford Linear Accelerator Center, Stanford University, Palo Alto, CA 94305, USA

- ^b Program in Applied Mathematics, Box 526, University of Colorado, Boulder, CO 80309, USA
- ^c Department of Physics and

Institute for Fusion Studies The University of Texas at Austin Austin TX 78712 USA

Received 25 January 1993 Revised manuscript received 12 April 1993 Accepted 14 April 1993

The effect of self-consistency on Hamiltonian systems with a large number of degrees of freedom is investigated

is observed that the system relaxes into a time asymptotic periodic state where only a few collective degrees are active;

low degree-of-freedom model is derived that treats the clump as a single *macroparticle*, interacting with the wave and chaotic sea. The uniform chaotic sea is modeled by a fluid waterbag, where the waterbag boundaries correspond

1-1 Introduction	first form non- in min officeto anal on "A mail dif	2.
Chaotic motion in Hamiltonian systems is	of degrees of freedom. Even in high dimensional	
gree of freedom [1]. Often, the systems studied are low dimensional approximations of many degree-of-freedom systems. In some cases, such	sional approximation, for example to study the motion of a single star in a given galactic grav- itational potential—this was the motivation for	
tions can be given with only a few degrees of freedom. However, there are many situations	erences in [1]). Such an approximation is not	
of degrees of freedom is essentially infinite. Generally, one expects such systems to exhibit greater chaos when the dimension increases;	in electromagnetic fields, where the fields pro- duced by the particles are ignored; the motion of tracer particles in a fluid, where the influence of these particles on the fluid velocity field is	
¹ Posthumous. Prepared by J.D.M. and P.J.M.	ignored (the passive advection problem); and	

0167-2789/94/\$07.00 (C) 1994 - Elsevier Science B.V. All rights reserved

the so-called "kinematic" dynamo, where a velogity field can intensify a magnetic field but

15 OUGA TOUCTION OF THE HEIR IS ISHOTOG. There has been little work on the effect of self-<u>ooncintanou In this paper we chow how it is pea</u> sible in a system with a large number of degrees of freedom for the inclusion of self-consistency. to result in dynamics with "effectively" few degrees of freedom.

to the formation of electrostatic plasma waves. Following [2] we suppose that the most unsta

010 01 111000 THAT 00 PROGOTITITICOD, 1110 11051001 1110 remainder of the modes-this is easily justified during the linear next of the avolution OWM showed that the wave grows in amplitude until it trans the beam particles. It then saturates and begins to oscillate in amplitude as the beam particles slosh in the wave potential. At this

and a second	<u>I am a second s</u>	
	Hamiltonian for each particle has one and a	We neglect these modes; this is justified, for
	half degrees of freedom, and so the motion	example, if the system has a finite length, and
	can be chaotic. However, each of the parti-	the sideband wavenumbers are forbidden by
	cles is charged and therefore contributes to the	periodic houndary conditions
	polential this is the sen-consistent encet. In	The usernations of the single wave after satura-
	of the field is given. Thus each particle experi-	a rigid bar in phase space. When the beam is
	extent that the other particles contribute to the	tate. Mynick and Kaufman computed the fre-
	single mode of the field. This is in contrast to	quency shift and amplitude oscillations of the
·	Ales Collinsolf consistent of he de demonstration of hereit	the second secon
<u> </u>		
<u></u>	particle. This latter case is considerably more	of the plasma wave oscillates, the beam parti-
	difficult.	cles can experience chaotic motion. They stud-
	Models similar to the one described above	ied the motion of a test particle in a model of
	may be appropriate for many physical situations;	this oscillating field and showed that much of the
	for example, a galaxy with a predominantly axi-	test particle phase space is indeed chaotic. How-
	symmetric gravitational potential that is per-	ever, there is an island in the phase space where
	turbed by a small humber of modes, say mose	the motion is regular, they noted that some nac-
.	corresponding to eniral density ways Each	tion of the barm norticles in the nur arian lar
۱	star contributes to these modes, and also has a	permittentes et e vivit should find memberves m
	possibly chaotic motion in the corresponding	the correct region of phase space to be trapped
	field. Similar effects occur for planetary rings,	in this oscillating island. Later Adam, Laval and
sti i	heam-heam interactions in accelerators tearing	Mendonce [7] studied a model in which a single
11	────▲ <u>─────↓ 1</u>	· ·
	O'Neil, Winfrey and Malmberg (hereafter re-	this two degree-of-freedom system is integrable
• 77		
	beam of charged particles moves in a back-	it was shown that the macroparticle system has
		and a firmer which are summarian difference in the

ground neutral plasma. The system is unstable

solutions which correspond to periodic oscilla-

tions of the bunch in the wave.

Related self-consistent problems include the interaction of a single particle with many waves [8] and the interaction of one wave with many other waves [9]. The more complicated

wave-particle turbulence, and it is not clear if the analysis of this paper can give any insight into this case.

In Section 2 we review the derivation of the OWM model, obtaining the Hamiltonian formu-

equations. Section 3 discusses numerical solutions of the OWM equations with up to 10⁵ beam

served at least 100 periods of these oscillations; as far as we can determine, the oscillations persist and the system becomes asymptotically pe-

ticle in this periodic potential, showing that a substantial portion of the original beam is indeed trapped in a stable island in the test par-

beam finds itself in the chaotic region of phase space, and spreads more or less uniformly over this region. The upper and lower boundaries of this "chaotic sea" are formed from invariant tori of the test particle system.

In Section 4, we construct a four degree of

wave, the second corresponds to the trapped

side the oscillating separatrix of the wave. We model these boundaries with sinusoidal curves, an assumption consistent with that of the single mode in the potential. Finally, the frequency shift of the trapped particle oscillations due to

2. Single wave model

O'Neil, Winfrey, and Malmberg (OWM) [2]

growth and saturation of the weak beam-plasma instability. In this section we briefly review the

sented by Mynick and Kaufman [5], discuss linear instability, and finally consider a special case where only a single beam particle is included

2.1. Derivation

To obtain the single wave model, the response

separately. We consider only the one dimensional, collisionless, nonrelativistic, electrostatic case. The total electron density

$$n(x,t) = n_{\rm p}(x,t) + n_{\rm b}(x,t)$$

IIIGai IUIGE equation

tions of the boundary of the chaotic sea and are derived from the "waterbag" approximation.

waterbog consist

timation. $m\ddot{x}_j = -eE(x_j, t),$

(2)

case the simulations show that the phase space density of the chaotic particles is indeed nearly constant and the boundaries of the chaotic zone are formed from invariant surfaces well outphase velocity of the resulting instability is much larger than the velocities of particles in the background plasma: the plasma responds nonresonantly, and trapping effects of plasma particles

	in the wave can be neplected. This implies that	At this point we accume that the electrostatic
1		
† Q		
-	$4\pi e n_{\rm p}(x,t) = (1 - \hat{\epsilon})\varphi''(x,t), \qquad (3)$	k-space is relatively narrow in units of $2\pi/L$.
	where φ is the electrostatic potential, $E = -\varphi'$. Substituting this into Poisson's equation for φ	In this case, if k represents the most unstable mode, the amplitude of all other Fourier com-
- ا		
	$\widehat{\epsilon}\varphi = 4\pi e n_{\rm b}(x,t) . \tag{4}$	single wave during the linear growth stage. Of course, some time after nonlinear saturation of
0		
	tion is most easily treated by Fourier transform.	stable spectrum depends on the small parameter $(n_{\rm b}/n_{\rm p})^{1/3}$, so that the single wave model will be
,	$\mathbf{e}_{\mathbf{r}} = \mathbf{e}_{\mathbf{r}} + $	most addrodriate in the weak deam case.
~	that the electrostatic response is dominated by	yields
-		

is a reasonable approximation to expand ϵ about one such zero retaining only the first derivative of ϵ with respect to ω :

$$\epsilon(k,\omega) \approx \epsilon(k,\omega_0) + \left. \frac{\partial \epsilon}{\partial \omega} \right|_{\omega=\omega_0} (\omega - \omega_0)$$

= $\epsilon'(\omega - \omega_0)$. (5)

For example, for a cold plasma $\epsilon = 1 - \omega_p^2/\omega^2$, and $\partial \epsilon / \partial \omega |_{\omega = \omega_0} \equiv \epsilon' = 2/\omega_p$. Transforming back to the time domain and using Eq. (4) then gives

$$\dot{E}_k + \mathrm{i}\omega_0 E_k = \frac{4\pi e}{kL\epsilon'} \sum_{j=1}^N \mathrm{e}^{-\mathrm{i}kx_j(t)}, \qquad (6)$$

beam density of Eq. (1):

$$= \frac{1}{L} \sum_{j=1}^{N} e^{-itw_{j}(x)}, \qquad (7)$$

$$\dot{p}_j = -e \left(E_k \, \mathrm{e}^{\mathrm{i}kx_j} + E_{-k} \, \mathrm{e}^{-\mathrm{i}kx_j} \right) \,.$$
 (8)

Equation (8) together with Eq. (6) are the closed dynamical system that governs the interaction of a single wave with the beam particles.

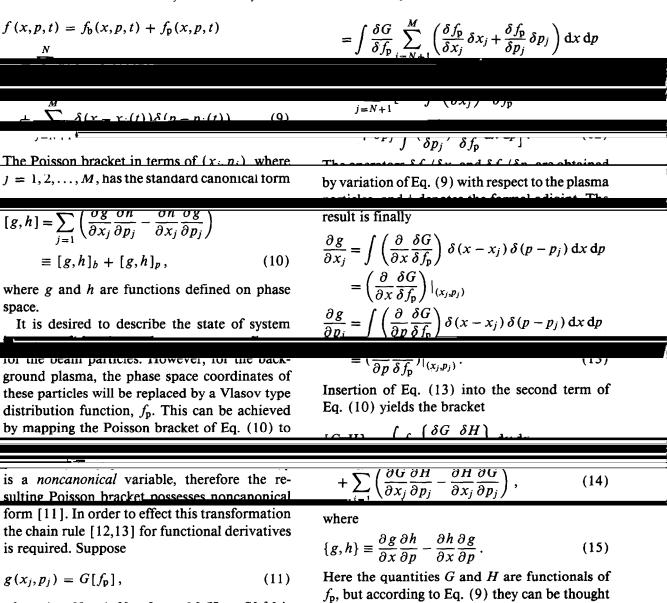
2.2. Hamiltonian structure and derivation

Now consider the derivation of the equations of motion, Eqs. (6) and (8), within the Hamiltonian context. The derivation proceeds by first considering the kinematics, i.e. the dynamical variables used to describe the state of the system, and then the dynamics obtained by finding the

appropriate Hamiltonian.

We begin the first part by supposing that the electrons are described by specifying their

1, 2, ..., M. The first N(< M) of these particles are singled out to represent the beam dynamics, while the remaining M - N particles represent the background plasma. The phase



where j = N + 1, N + 2, ..., M. Here $G[f_p]$ is a functional of f_p ; the relationship between the phase space function g and the functional G is

is obtained by varying both sides of this equation:

$$\delta g = \sum_{j=N+1}^{M} \left(\frac{\partial g}{\partial x_j} \, \delta x_j + \frac{\partial g}{\partial p_j} \, \delta p_j \right)$$
$$= \delta G$$

 $g(x_i, p_i) = G[f_{\mathbf{p}}],$

space.

ground Vlasov plasma electrons is obtained by inserting

of as ordinary functions of the beam particle co-

ordinates (x_j, p_j) where j = 1, 2, ..., N. Note

that discreteness has now disappeared from f_p .

$$f(x, p, t) = f_{p}(x, p, t) + \sum_{j=1}^{N} \delta(x - x_{j}(t)) \delta(p - p_{j}(t))$$
(16)

4177

$H[f_{\mathbf{p}}; x_{j}, p_{j}] = \frac{1}{2m} \int p^{2} f_{\mathbf{p}} \, \mathrm{d}x \, \mathrm{d}p - \frac{e}{2} \int f_{\mathbf{p}} \varphi_{\mathbf{p}} \, \mathrm{d}x + \sum_{i=1}^{N} \left(\frac{p_{j}^{2}}{2m} - e \varphi_{\mathbf{p}}(x_{j}) - \frac{1}{2} e \varphi_{\mathbf{b}}(x_{j}) \right), \quad (18)$

where $\varphi_p(x_j)$ and $\varphi_b(x_j)$ are the contributions to the electrostatic potential of the plasma and

yields the hybrid system.

Now we can turthtouthe task of obtaining, from the hybrid system, the approximate system of

sumed to be described by an equilibrium distribution function of compact support in velocity, plus the single linear wave, whose phase velocity

wave-barticle effects are eliminated in the back-

bation of the distribution function, the analysis of [14] and [15] implies that the linearization of the plasma energy becomes identically the wellknown expression for the dielectric energy of a plasma wave. Second, the self-interaction potential of the beam, φ_b , is neglected in comparison to that of the plasma, φ_p , a justifiable assumption in light of smallness of n_b/n_p . Thus, Eq. (18) becomes

$$H(E_{k}, E_{-k}, x_{j}, p_{j}) = \frac{L}{4\pi} \omega_{0} \epsilon' |E_{k}|^{2} + \sum_{j=1}^{N} \left(\frac{p_{j}^{2}}{2m} - \frac{i}{k} e_{k} e^{ikx_{j}} + \frac{i}{k} e_{-k} e^{-ikx_{j}} \right).$$
(19)

It remains to find the appropriate Poisson bracket in terms of E_k and E_{-k} instead of f_p . Since the plasma is in essence being modeled as a fluid, an easy way to obtain this is to map

$$[g,h] = \sum_{j=1}^{N} \left(\frac{\partial g}{\partial x_j} \frac{\partial h}{\partial p_j} - \frac{\partial h}{\partial x_j} \frac{\partial g}{\partial p_j} \right) - \frac{i}{L} \frac{4\pi}{\epsilon'} \left(\frac{\partial g}{\partial E_k} \frac{\partial h}{\partial E_{-k}} - \frac{\partial h}{\partial E_k} \frac{\partial g}{\partial E_{-k}} \right).$$
(20)

Eqs. (6) and (8) in the form

The bracket of Eq. (20) is not quite canonical; however, with the substitution

$$E_{-k} = i \left(\frac{4\pi}{L\epsilon'}\right)^{1/2} \mathcal{J}^{1/2} e^{i\vartheta}, \qquad (22)$$

the electric field is expressed in terms of conven-

tional action-angle variables, and

$$[g,h] = \sum_{j=1}^{N} \left(\frac{\partial g}{\partial x_j} \frac{\partial h}{\partial p_j} - \frac{\partial h}{\partial x_j} \frac{\partial g}{\partial p_j} \right) + \left(\frac{\partial g}{\partial \vartheta} \frac{\partial h}{\partial \mathcal{I}} - \frac{\partial h}{\partial \vartheta} \frac{\partial g}{\partial \mathcal{I}} \right), \quad (23)$$

while the Hamiltonian of Eq. (19) becomes

$$H(\vartheta, \mathcal{J}, x_j, p_j) = \omega_0 \mathcal{J} + \sum_{j=1}^N \left[\frac{p_j^2}{2m} - \frac{2e}{k} \left(\frac{4\pi}{L\epsilon'} \right)^{1/2} \mathcal{J}^{1/2} \cos(kx_j - \vartheta) \right].$$
(24)

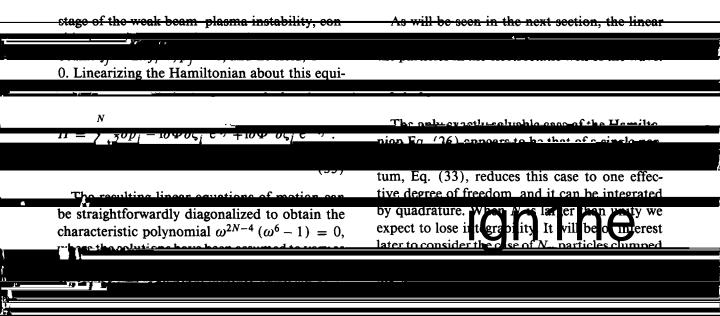
To complete the derivation, it is convenient to introduce scaled, dimensionless variables based on the fundamental frequency,

$$\omega_{\rm b}^3 = \frac{4\pi \ {\rm e}^2 N}{mL\epsilon'}.\tag{25}$$

Here ω_b is a harmonic mean of the beam's plasma frequency and $1/\epsilon'$, which is of order

$$\frac{d\Phi}{d\tau} = [\Phi, H] = [\Phi, \Phi^*] \frac{d\Phi}{\partial \Phi^*}, \quad (31)$$

$$H(J, \theta, p_J, \xi_J) = \sum_{j=1}^{N} \left[\frac{1}{2} p_J^2 - 2 \left(\frac{J}{N} \right)^{1/2} \cos(\xi_J - \theta) \right], \quad (26)$$
Where the dimensionless variables are defined by
$$\frac{d\Phi}{d\tau} = \frac{i}{N} \sum_{j=1}^{N} e^{-i\xi_J}, \quad (32)$$
Note that these equations hold for arbitrary choices of the physical parameters, such as e/m such as e/m such as $h = 1$. (32)
Note that these equations hold for arbitrary choices of the physical parameters, such as e/m such asuch as e/m such asu



the flow (recall that if ω is an eigenvalue then ω^* , $-\omega$ and $-\omega^*$ must also be eigenvalues). In dimensional units, using Eq. (27), we have

$$\hat{\omega}_j = \omega_b e^{ij\pi/3}, \quad j = 0, 1, \dots, 5,$$
 (36)

which includes the unstable beam-plasma mode (the case j = 2). We can physically identify the eigenmodes by considering the equations of motion. Differentiating the equation for Φ twice and substituting for ξ gives

$$\frac{c^2 c \epsilon}{d\tau^3} = \frac{1}{N} \sum_{j=1}^N e^{-\kappa_j} \,\delta\xi_j = i\delta\Phi\,, \qquad (37)$$

upon noting that $\sum e^{-2i\xi_j^0} = 0$. This shows that the nonzero frequencies are associated with nonzero Φ . The eigenmodes for the conjugate roots, ω^* , $-\omega$ and $-\omega^*$, are the same as that for ω except for varying choices of signs. The remaining roots of the characteristic equation $(\omega = 0 \text{ of multiplicity } 2N-4)$ have eigenmodes

N-2 independent solutions of $\sum e^{-\kappa_j} \delta \xi_j = 0$. The double multiplicity of each of these roots

$$H = \frac{p^2}{2N_{\rm m}} - 2N_{\rm m} \left(\frac{J}{N}\right)^{1/2} \cos\left(\xi - \theta\right), \qquad (38)$$

where $p_{\overline{\text{cash}}}N_{m}p_{1} = N_{m}p_{2}...$ is the macroparticle momentum. The Hamiltonian *H* can be reduced to one degree of freedom by defining the total momentum P = p + J as before to obtain

$$H = \frac{p^2}{2N_{\rm m}} - 2N_{\rm m} \left(\frac{P-p}{N}\right)^{1/2} \cos\psi.$$
 (39)

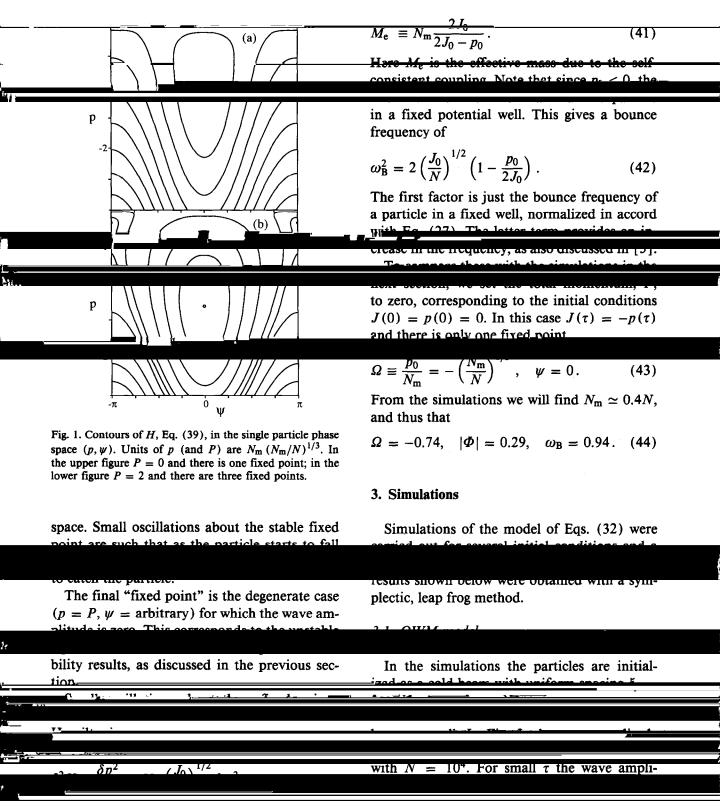
The equations for this case were studied in detail by Adam, Laval and Mendonca [7], who did not use the Hamiltonian approach.

nondegenerate fixed points. These occur at the points defined by

$$p_0^3 - p_0^2 P + N_m^3 \frac{N_m}{N} = 0, \quad \psi_0 = 0 \text{ or } \pi.$$
 (40)

The fixed point with $(p_0 < 0, \psi_0 = 0)$ is stable and corresponds to the macroparticle sitting in the bottom of the potential well. The two fixed points with $(p_0 > 0, \psi = \pi)$ are less intuitive. These exist only if $P > 3N_m (N_m/4N)^{1/3}$. They

potential well. The lower momentum particle is unstable, while the larger momentum particle is



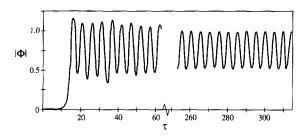


Fig. 2. Plot of $|\Phi(\tau)|$, the normalized wave amplitude, for N = 10000 particles initialized as a cold beam.

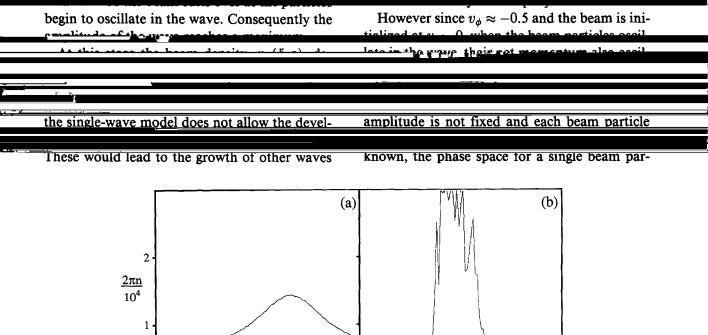
predicted by Eq. (36) with the phase velocity $v_{\phi} = \mathcal{R}(e^{2\pi i/3}) = -0.5$. As the wave grows, the beam experiences a growing sinusodal perturbation, and as can be seen in the density plot of Fig. 2-th p-page density also veries sinussidally.

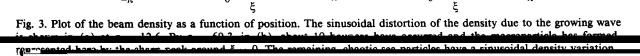
-π

and undoubtedly greatly change the subsequent behavior of the system.

None-the-less, the subsequent development of the OWM dynamics is quite interesting. As the beam particles begin to oscillate in the wave, their oscillation frequencies depend upon their energy, just as for a single particle in a fixed potential. Thus as the beam begins to rotate about the potential minimum, those particles closer to the center have larger oscillation frequencies than those near the "separatrix".

If the wave amplitude were fixed, one would see phase mixing of the particles (visualized as an ever tighter spiral in the particle phase space), and the oscillations in the particle total energy would damp away this is the mechanism of Landau damping in a large emplitude





 $-\pi$

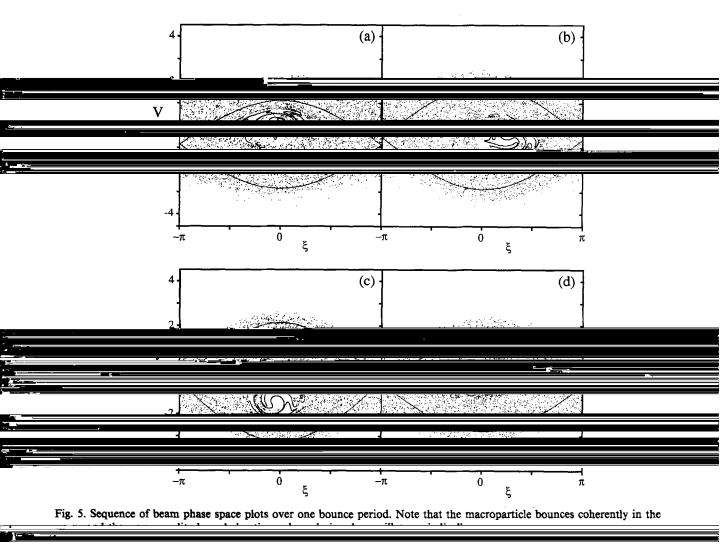
		"chaotic sea" The remaining 40% of the parti-
		wave as seen in the sequence of phase space
، ريد		instantaneous somerstrip of the wave that is
ـــــــــــــــــــــــــــــــــــــ		
	0 π 2π	of the cluster oscillate 180° out of phase. This
	Fig. 4. Plot of the beam particle phase space at $\tau = 641$ showing a well defined macroparticle and chectic sectors	the macroparticle, in the model of Section 4. In addition to the uniform cold beam, several
•	tisle in a since as illeting and at the local	different initial conditions have been partially
	tohere47	$\frac{1}{1}$
	· · · · · ·	
<u>.</u>	of the beam particles at a fixed time. Note the	these cases some subset of the particles remained
	with a nearly uniform density, and the other a	system did not appear to settle into an equilib-
	more coherent cluster of particles. In the cluster	rium. Cold beams with nonzero momenta also
	one still sees evidence of the initial beam though	lead to oscillations as was shown in [2] though
· ·	In the simulations which were carried out up	the momentum We have not investigated this in.
	sisted, and indeed, as can be seen in Fig. 2 the	initial conditions that will give rise to a periodic
	oscillations become increasingly periodic with	initial conditions that will give rise to a periodic final state.
	oscillations become increasingly periodic with time. Furthermore as we varied N up to 10^5 and	final state.
	oscillations become increasingly periodic with time. Furthermore as we varied N up to 10^5 and improved the integration accuracy, we noticed that these oscillations became more periodic and	
	oscillations become increasingly periodic with time. Furthermore as we varied N up to 10^5 and improved the integration accuracy, we noticed that these oscillations became more periodic and constant in amplitude as the number of particles	final state.
	oscillations become increasingly periodic with time. Furthermore as we varied N up to 10^5 and improved the integration accuracy, we noticed that these oscillations became more periodic and constant in amplitude as the number of particles increased and as the accuracy improved. Thus	final state. 3.2. Test particle To investigate further dynamics of the beam particles, consider the "test particle" motion of
	oscillations become increasingly periodic with time. Furthermore as we varied N up to 10^5 and improved the integration accuracy, we noticed that these oscillations became more periodic and constant in amplitude as the number of particles	final state. 3.2. Test particle To investigate further dynamics of the beam particles, consider the "test particle" motion of a single particle in a given oscillating potential.
	oscillations become increasingly periodic with time. Furthermore as we varied N up to 10^5 and improved the integration accuracy, we noticed that these oscillations became more periodic and constant in amplitude as the number of particles increased and as the accuracy improved. Thus we believe that the asymptotic state is a periodic	final state. 3.2. Test particle To investigate further dynamics of the beam particles, consider the "test particle" motion of

particles—those with relatively large energies in the wave frame—experience chaotic motion, and spread out roughly uniformly in a region of phase space whose average width is $\Delta \omega = 4.7$.

where I and A are considered to be given nori

(45)

 $H_{\rm t}(p,\xi,\tau) = \frac{1}{2}p^2 - 2\left(\frac{J(\tau)}{N}\right)^{1/2} \! \cos\left(\xi\!-\!\theta(\tau)\right), \label{eq:holescaled}$



Here we determine J and θ numerically, from the simulations of Section 3.1, building these functions from an average over a number of pe-

A stroboscopic plot of the test particle dynamics is shown in Fig. 6 for several different values of θ . The dots represent the trajectories of a number of different test particles. As was also noted in [6], there is a prominent stable island in the test particle phase space which oscillates exactly out of phase with the potential; much of the rest of the phase space is chaotic. Also shown in the plots represents the position of one of the 10000 beam particles. Note that the macroparticle clump site as pear as can be ascertained.

verifies an assertion in [6], where it was merely noted that some fraction of the beam particles initially stretched across the position of the test particle island.

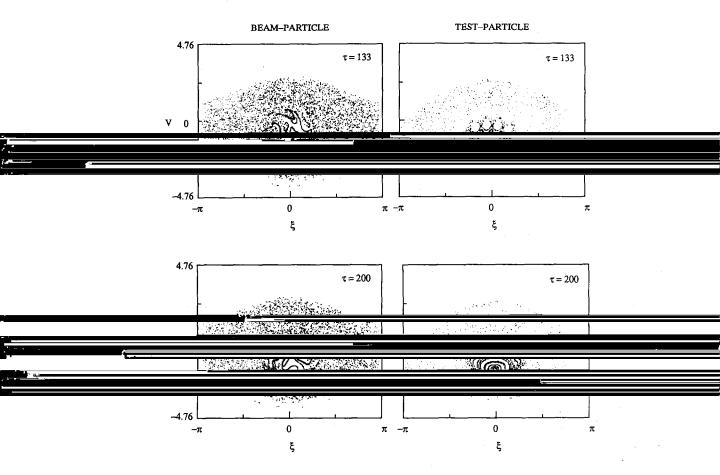


Fig. 6. Phase space plots comparing the full simulation of Section 3 with the dynamics of a test particle in a given time-dependent potential $\Phi(\tau)$ as determined by the simulation. Shown are several test particle initial conditions at three different times during a cycle.

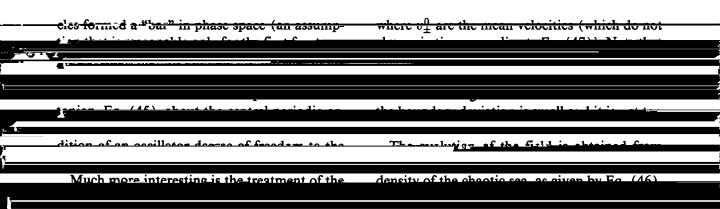
4. Chaotic sea model

Aso we have set of the simulations, the asymptotic state of the cold beam initial condition, evolved under the OWM Hamiltonian, appears to be almost exactly periodic. Approxinstals 10% of the initial have forme a clumm of particles that oscillates in the potential well formed by the wave. The remaining particles freedom, which approximately describes the full 10 001 degree-of-freedom system. In the model, as noted above, we assume that the clump of regularly oscillating particles is localized enough so that all these particles can be treated as one located at (ξ, p) . This macroparticle contains $\frac{N}{restrictes}$ and hence a more $\frac{N}{restrictes}$ and hence $\frac{N}{restrictes}$ and $\frac{N}{restrictes}$ are $\frac{N}{restrictes}$ and $\frac{N}{restrictes}$ and $\frac{N}{restrictes}$ and $\frac{N}{restrictes}$ and $\frac{N}{restrictes}$ and $\frac{N}{restrictes}$ are $\frac{N}{restrictes}$ and $\frac{N}{restrictes}$ and $\frac{N}{restrictes}$ and $\frac{N}{restrictes}$ are $\frac{N}{restrictes}$ and $\frac{N}{restrictes}$ are $\frac{N}{restrictes}$ and $\frac{N}{restrictes}$ and $\frac{N}{restrictes}$ are $\frac{N}{r$

can be treated as a single particle ignores any in-

reduced Hamiltonian model of four degrees of

man [5] who assumed that the cluster of parti-



phase space density of these particles appears to

In velocity space. We assume that these particles can be treated as a continuum with a constant phase space density f between the boundaries

$$\frac{4\pi c}{k\epsilon'} \left(f_{\rm c}(\tilde{v}_+ - \tilde{v}_-) + \frac{4\pi}{L} e^{-ikx_m} \right).$$
 (50)

These acceptions are non-dimensionalized us

(51)

$$n_c(x,t) = \int_{v_-}^{v_+} f_c \, \mathrm{d}v = f_c(v_+ - v_-) \tag{46}$$

ber of such particles in the length L will be de-

noted by $N_c = N - N_m$. Particles in the chaotic sea evolve according to Eq. (2), and hence f_c evolves according to the Vlasov equation. As is

two equations for the evolution of the boundaries [18]. These equations are called the wa-

$$\frac{\partial v_{+}}{\partial t} + v_{+} \frac{\partial v_{+}}{\partial x} = -eE,$$

$$\frac{\partial v_{-}}{\partial t} + v_{-} \frac{\partial v_{-}}{\partial x} = -eE.$$
 (47)

Following the philosophy of the derivation of

$$\begin{split} \dot{\boldsymbol{\Phi}} &= \mathrm{i} \frac{N_{\mathrm{c}}}{N \Delta \omega} \left(V_{+} - V_{-} \right) + \mathrm{i} \frac{N_{\mathrm{m}}}{N} \, \mathrm{e}^{-\mathrm{i}\xi} \,, \\ \ddot{\boldsymbol{\xi}} &= \mathrm{i} \boldsymbol{\Phi} \, \mathrm{e}^{\mathrm{i}\xi} - \mathrm{i} \boldsymbol{\Phi}^{*} \, \mathrm{e}^{-\mathrm{i}\xi} \,, \end{split}$$

In terms of these variables the equations of mo-

 $\omega_{\pm} \equiv \frac{k v_{\pm}^0 - \omega_0}{\omega_{\rm b}} \,, \quad V_{\pm} \equiv \frac{k}{\omega_{\rm b}} \widetilde{v}_{\pm} \; {\rm e}^{{\rm i}\omega_0 t} \,.$

where $\Delta \omega = \omega_+ - \omega_-$ is the average, nondimensional width of the chaotic sea.

This set of equations is also a Hamiltonian system, with the wave action-amplitude variables defined in Eq. (28), and the new actionamplitude variables for the chaotic sea defined by

$$V_{+} = \left(\frac{J_{+}\Delta\omega}{M}\right)^{1/2} e^{-i\theta_{+}},$$

$\ln v_{\pm}(x)$:

terbag equations:

$$v_{\pm} = v_{\pm}^0 + \widetilde{v}_{\pm} e^{ikx} + \widetilde{v}_{\pm}^* e^{-ikx} . \qquad (48)$$

The equations of motion then become

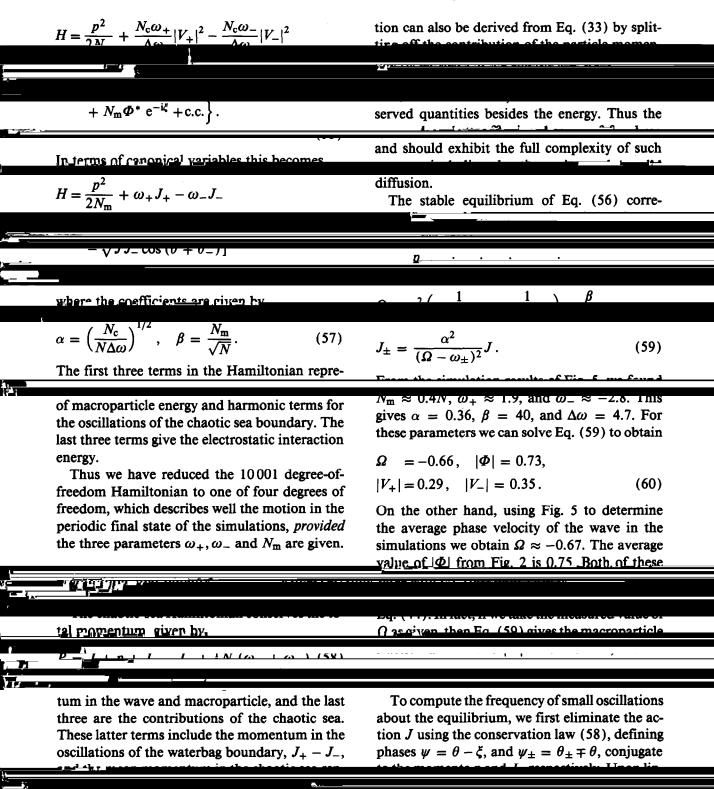
$$\left(\frac{\partial}{\partial t} + ikv_{\pm}^{0}\right)\tilde{v}_{\pm} = -eE_k, \qquad (49)$$

which results in the Poisson bracket relations

$$[V_{\pm}^*, V_{\pm}] = \pm i \frac{\Delta \omega}{N_c}.$$
 (54)

The Hamiltonian takes the form

 (N_c)



$- \delta^{2} H_{\overline{\tau_{1}}} + \delta \mathbf{n} \cdot \mathbf{M}^{-1} \cdot \delta \mathbf{n} + \frac{1}{2} \delta \mathbf{w} \cdot \mathbf{K} \cdot \delta \mathbf{w} \qquad ($	61) much better than the single particle calculation
$\delta \psi_+, \delta \psi$) are the deviations from equilibri	um.
The most in B.f. ale	
tive definite. The matrix K, the effective spi	fing
constant matrix, turns out to be diagonal. In	
der mat me unee terms m & be positive, we n	electrosistic interscition of many particles with a
assume that $\psi = \psi_+ = 0$, while $\psi = \pi$.	nlasma wave. The wave arises from an instabil-
is consistent with the fact that the lower bou	ity (the beam-plasma instability) of the initial
ary of the chaotic sea is observed to have a 1	state corresponding to a cold hear of particles
phase lag with respect to the upper boundar	y. In the simulations, the asymptotic state corre-
The frequencies of small oscillation are gi	sponds to a periodically oscillating wave ampli-
by the square roots of eigenvalues of the ma	The together with a trapped clump of particles
KM^{-1} . For the parameters of the simulation,	
mass matrix is diagonal to a good approxi	
tion. The element M_{11}^{-1} turns out to be ident to $1/M_e$ of Eq. (41); neglecting terms of or	
J_{\pm}/J , the other diagonal elements are	chaotically—becoming successively trapped and
J_{\pm}/J , the other diagonal elements are	detrapped.
$M^{-1} - \frac{1}{2} \frac{\Omega - \omega_{+}}{\omega_{+}} = M^{-1} - \frac{1}{2} \frac{\Omega - \omega_{-}}{\omega_{-}}$	We modelled this motion by a four degree-of-
	chaotic sea, and one to the wave. This model
The matrix K is	quantitatively captures the asymptotic state of
	the effectively infinite degree of freedom and
Using the values obtained before for the ed	- One would mae to speculate mat mere are
librium, we determined the eigenvalues num	
ically from the full matrix. The three oscillat	
frequencies are	ple in the case of galactic dynamics, the self-
	Appistent pronection of a density wave would
A NO MORE OF MARKE OF MARKE OF	
the frequency of oscillations of the summer	atian the OWM model What is the "hasin" of initial
The tradilesses of cooling on the course	pha the limit model what is the "basin" of initial
mant with the coloulated value We have no	ari for momente consider a momente bacministic state
	ter- discussed at the end of Section 3.1.
modes; however, it might be possible to de	
modes; however, it might be possible to de mine these through careful simulation.	- Is there a way of self-consistently calculating

16

	II Tommer at 1 Call consistent	along the deal of the second s
partie [<u>s_the</u>	ere a periodic state of the many particle sys-	nonlinear deam-plasma interaction, Phys. Fluids 21 (1978) 653-663
	the ONUS model, only a simple Fourier har	motion in a large amplitude plasma wave, Phys. Fluids
the ef	ffect of adding additional harmonics?	nonlinear Langmuir waves, Phys. Fluids 24 (1981) 260–267.
the fu	ull array of possible Hamiltonian motions.	ear diffusion far from the chaotic threshold, Physical <u>Bayleyy Letters 65 (1990) 3132-3135</u>
mani	tolds leading to homoclinic phenomena.	three-wave interaction: integrable case of this system
states	s that correspond to these motions?	 [10] J.R. Cary, I. Doxas, D.F. Escande and A.D. Verga, Enhancement of the velocity diffusion in longitudinal plasma turbulence, Phys. Fluids B 4 (1992) 2062. [11] P.J. Morrison, The Maxwell-Vlasov equations as
Ackne	owledgements	a continuous Hamiltonian system, Phys. Lett. A80 (1980) 383-386.
the at	utitors were at the institute for Fusion Stud-	Princeton Plasma Physics Laboratory Report PPPL- 1783 (1981), available as American Institute of Physics
was p	provided by the US Dept. of Energy Con- No. DE-FG05-80ET-53088 to the Univer-	Publication Service, 335 East 45th Street, New York, NY 10017). [13] A.N. Kaufman and R.L. Dewar, Canonical derivation
<u>sity_o</u>	of Terre at Aristin	of the Vlasov-Coulomb noncanonical Poisson bracket,
Refer	rences	 [14] I.J. Morrison and D. Thisen, Directine energy versus plasma energy, and Hamiltonian action-angle variables for the Vlasov equation, Phys. Fluids B4 (1992) 3038- 3057. [15] P. Shedwick and P.L. Marrison, On-control plasma
S.	Nietenge 2 Deprint Selection (Adem Hilser 1097)	[16] C. Kueny, Ph.D. thesis Nonlinear instability and chaos
[2]]	plasma, Phys. Fluids 14 (1971) 1204–1212.	[17] T.M. O'Neil, Collisionless damping of nonlinear plasma oscillations, Phys. Fluids 8 (1965), 2255, 2262
	\mathbf{v} One ventering \mathbf{v}_{i} and \mathbf{v}_{i}	(Academic New York 1972) pp 45-47

monoenergetic electron beam, Plasma Phys. 14 (1972) 591–600.

(Academic, New York, 1972), pp. 45-47.